

# Nuclear Mass Predictions with Bayesian Neural Network

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# Introduction : Nuclear Mass

Theories of nuclear mass

## Macroscopic

LDM

## Microscopic

Hartree-Fock-Bogoliubov (HFB)

## Macro-Micro

FRDM

Duflo-Zuker (DZ)

Moller-Nix (MN)

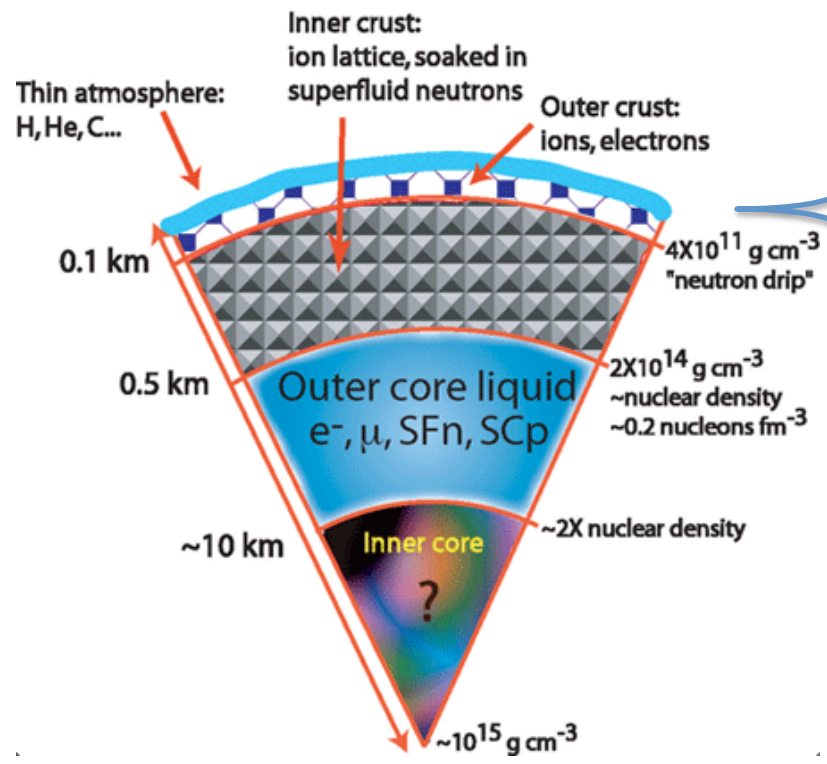
Experiments of nuclear mass

## Atomic mass evaluations (AME)

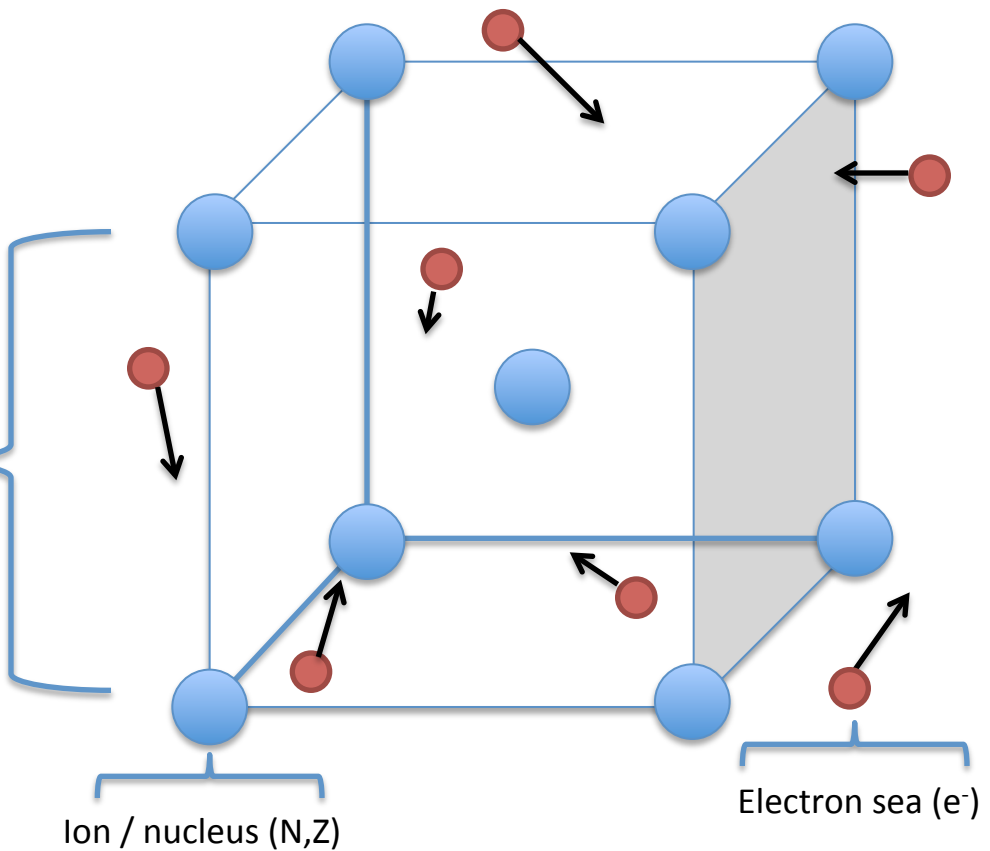
1988, 1993, 1995, 2003, **2012**

# Motivation : Why Mass?

## Neutron star outer crust



Composition

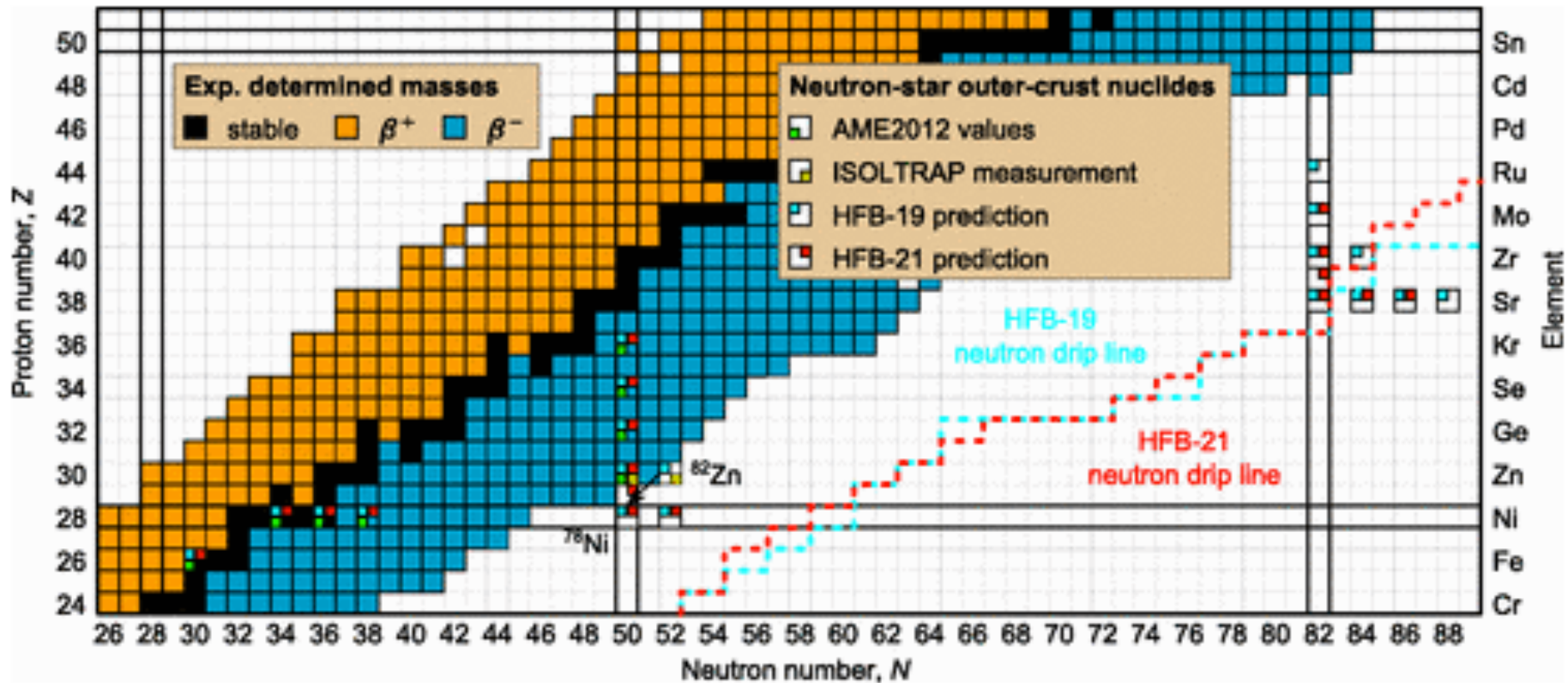


**Energy = Mass + Degeneracy + Lattice**

← (blue arrow)      → (red arrow with ?)

# Motivation : Crustal Nuclei

Outer crust by HFB



The constituents in multicolor

# Frameworks : Bayes' Theorem

Bayes' probabilities

$$\text{Posterior} \rightarrow P(A | B) = \frac{\text{Likelihood} \cdot P(A)}{P(B)} \leftarrow \text{Prior}$$

Example : Ebola Test

$$p(E) = 0.01$$

$$p(P|E) = 0.9$$

$$p(E|P) = \frac{p(P|E)p(E)}{p(P)}$$

$$p(E|P) = \frac{0.9 \times 0.01}{(0.9 \times 0.01) + (0.1 \times 0.99)} = 0.083$$

P -> positive ; E -> Ebola



# Frameworks : Neural network

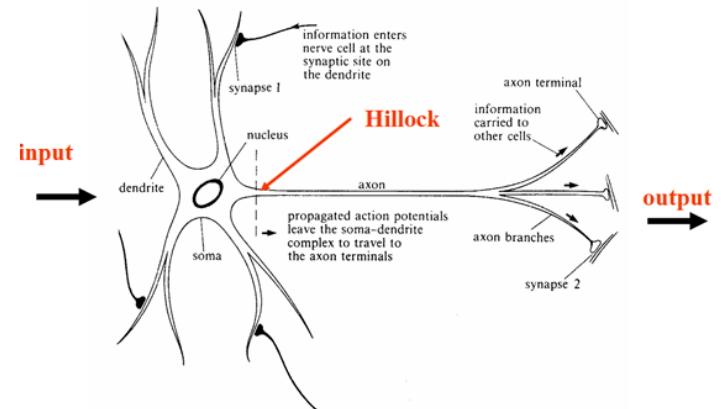
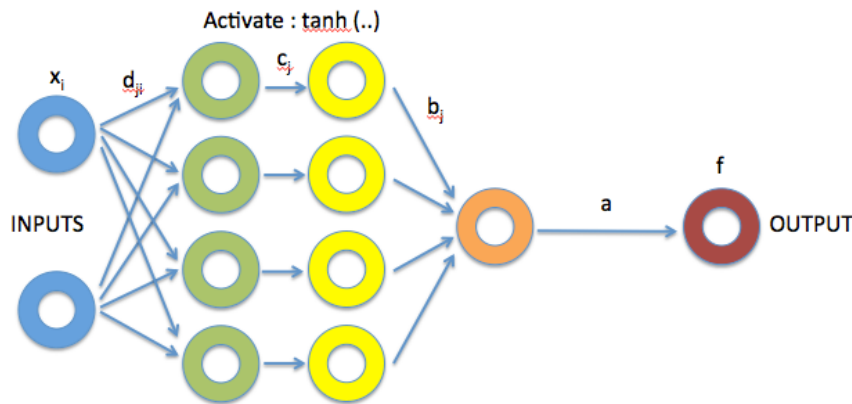
## Neural Network

### Basis function

$$f(\bar{x}, \bar{\omega}) = a + \sum_{j=1}^H b_j \tanh(c_j + \sum_{i=1}^I d_{ji} x_i)$$

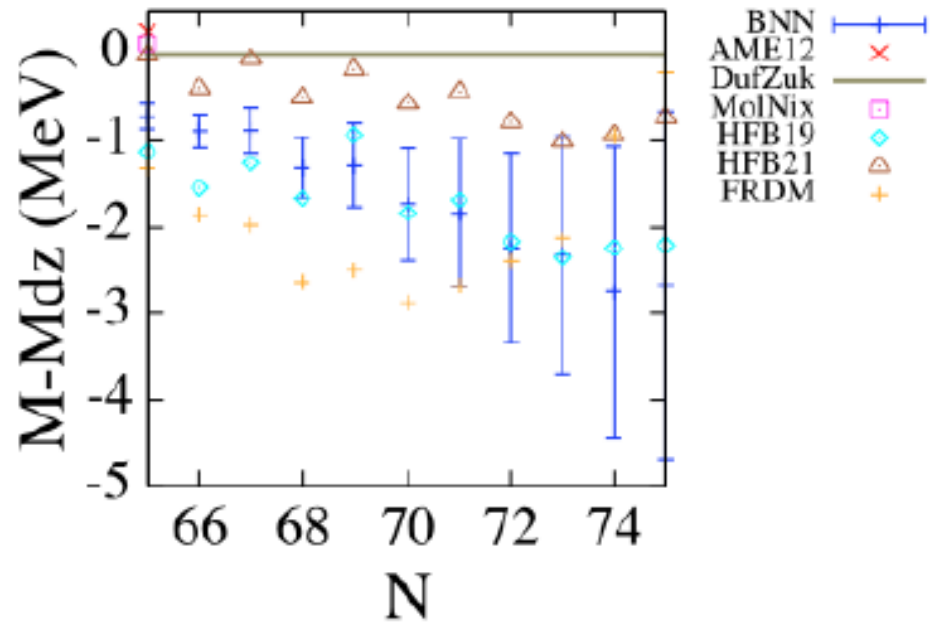
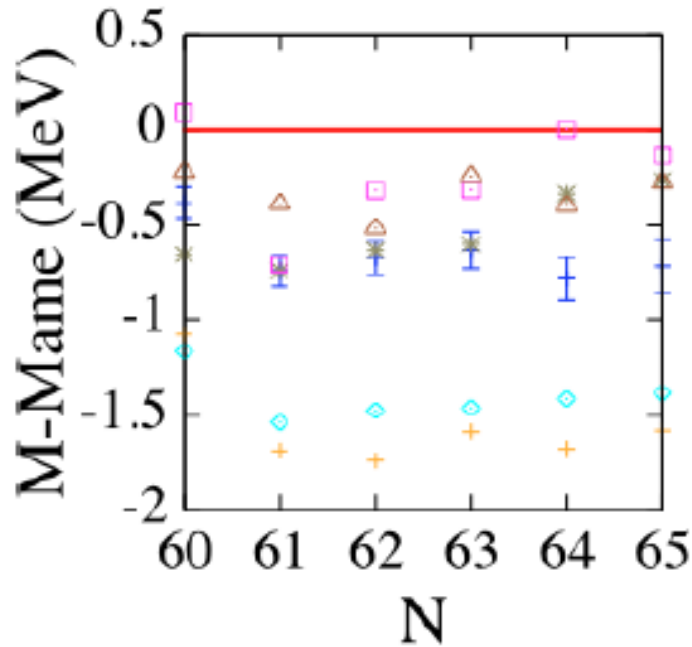
$$f(\bar{x}, \bar{\omega}) \rightarrow M(N, Z)$$

$x_i \rightarrow (N, Z)$



# Results : BNN Predictions

LDM with BNN on Krypton isotopes



At outer crust region :

$20 < N < 84$  and  $20 < Z < 50$

Model	LDM	DZ	MN	FRDM	HFB19	HFB21
$\sigma_{rms}^{(pre)}$ (MeV)	3.233	0.528	1.006	0.848	0.743	0.639
$\sigma_{rms}^{(post)}$ (MeV)	0.532	0.303	0.479	0.432	0.412	0.398
$\Delta\sigma / \sigma_{rms}^{(pre)}$	0.84	0.43	0.52	0.49	0.45	0.38

# Outlook

- Significance of nuclear mass prediction in Astrophysics :  
Outer crust of neutron star  
r-process
- Bayesian approach for data fitting
- Neural network as basis for any smooth function
- Potential use of BNN for many other nuclear observables



**Thank You**